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and $r =$ radius of the orbit, $m_s =$ mass of the satellite, $m_e =$ mass of the earth, $h_0 =$ orbit altitude, that is, the height above the subsatellite point on the earth terminal (m). The gravitational force F_g acting on the satellite of mass m_s at distance r from the center of the earth is $F_g = m_s g \frac{R_e^2}{r^2}$ (2.8) where $g =$ acceleration due to gravity at the surface of the earth $= 9.807 \text{ m/s}^2$, and $R_e =$ radius of the earth. The value varies with location. For example, R_e at the equator $= 6378.39 \text{ km}$ ($\approx 6378 \text{ km}$) R_e at the pole $= 6356.91 \text{ km}$ ($\approx 6357 \text{ km}$) K21636 Book.indb 27 10/22/13 4:09 PM 28 Satellite Communication Engineering, Second Edition

Consequently, for a satellite in a stable circular orbit round the earth, $F = F_g$ (2.9) In view of (2.7) and (2.8) in (2.9), $r^3 = g R_e^2 \omega^2$ (2.10) The period of the orbit, t_s , that is, the time taken for one complete revolution (360° or 2π radians), can be expressed as $t_s = \frac{2\pi r}{v} = \frac{2\pi R_e}{r^2} \frac{r^3}{g \omega^2}$ (2.11) If we assume a spherical homogeneous earth, a satellite will have an orbital velocity represented by $v = R_e \omega$ (2.12) For elliptical orbits, Equations (2.11) and (2.12) are also valid by equating the ellipse semimajor axis a with the orbit radius, r (i.e., $r = a$). In terms of the orbit parameters, r is replaced with the average of the apogee to the focus and the perigee to the focus. By definition, a perigee is the lowest altitude point of the orbit, whereas an apogee is the highest altitude point of the orbit. In a circular orbit, with variable altitude and upon substitution of empirical values in (2.11) and (2.12), Figure 2.4, which relates period T (h) 25 7 Time 20 6 5 15 4 10 3 Velocity v (km/s) 2 5 0 1 0 10 20 30 Orbit Altitude $\times 10^3$ (km) 40 Orbital Velocity (km/s) 9 0 Figure 2.4 Satellite period and orbital speed vs. altitude. K21636 Book.indb 28 10/22/13 4:09 PM Satellites 29 and velocity for circular orbits, assuming a spherical homogeneous earth, is plotted. For a circular orbit at an altitude of 35,784 km, Figure 2.4 shows that a geosynchronous orbit takes a period of rotation of the earth relative to the fixed star (called sidereal day) in 86,163.9001 s, or 23 h 56 min 4 s. In some books and papers, an approximate value of 36,000 km is frequently cited for the altitude of the satellite in geosynchronous orbit. The geosynchronous orbit in the equatorial plane is called the geostationary orbit. Although a satellite in the geostationary orbit does not appear to be moving when seen from the earth, its velocity, from Figure 2.4, in space equals 3.076 km/s (11,071.9 km/h). Low-altitude satellites, which have orbits of less than nominally 24 h, have other applications in addition to those earlier tabulated in Table 2.1. The applications include reconnaissance purposes, provision for communications at extreme north and south latitudes when in a polar orbit, and numerous business opportunities in producing remotely located monitoring and data acquisition devices that could be accessed by satellite. 2.2.1 Orbital Errors It is not possible to put a satellite into a perfect geostationary orbit because any practical orbit is slightly inclined to the equatorial plane. In addition, it is not exactly circular; it does not have exactly the same period as that of the earth's rotation, and it is constantly bombarded by disturbing forces (such as the attraction of the sun and moon) that try to change the orbit. These disturbing forces cause satellites to drift slowly in longitude. Their effects are counteracted from time to time by operating thrusters on the satellite. It is logical to suggest that these forces introduce orbital errors outside the intended nominal longitudes. Advances in space technology have enabled minimization of the orbital errors. For example, INTELSAT V satellites were kept within $\pm 0.1^\circ$ of the equator and of their nominal longitudes, whereas INTELSAT VI satellites are kept within $\pm 0.02^\circ$ of the equator and $\pm 0.06^\circ$ of their nominal longitudes. 2.3 Coverage Area and Satellite Networks 2.3.1 Geometric Coverage Area Clarke [2] foresaw in his article that it would be possible to provide complete radio coverage of the world from just three satellites, provided they could be precisely placed in geosynchronous orbit. Figure 2.5 demonstrates this. K21636 Book.indb 29 10/22/13 4:09 PM 30 Satellite Communication Engineering, Second Edition Geosynchronous orbit Satellite Earth Figure 2.5 Complete coverage of the earth's surface from three satellites. The amount of coverage is an important feature in the design of earth observation satellites. Coverage depends on altitude and look angles of the equipment, among several factors. To establish the geometric relationship of the coverage, we take a section of the satellites in Figure 2.5 as an illustration. This section is redrawn as shown in Figure 2.6. The maximum geometric coverage can then be defined as the portion of the earth within a cone of the satellite at its apex, which is tangential to the earth's surface. Consider the angle of view from the satellite to the earth terminal as α ; then the apex angle is 2α . The view angle has a mathematical physical function given by $\frac{R_e}{r} \cos \alpha = \sin^{-1} \left[\frac{R_e}{r} \right] = \sin^{-1} \left[\frac{R_e}{R_e + h_0} \right]$ (2.13) Using empirical values ($R_e = 6378 \text{ km}$, $r = 42,162 \text{ km}$), the apex angle 2α equals 17.33° , the planar angle beamwidth. It follows that an "earth coverage" satellite antenna must have a minimum beamwidth θ_{BW} of 17.33° . In practice, an antenna of 18° or 19° beamwidth is used to allow for directional misalignment. Thus, for a single geostationary satellite to illuminate in excess of a third of the earth's surface, the antenna minimum beamwidth must be at least 2α . The beamwidth of the satellite antennas determines the area of the earth serviced or covered. The beamwidth required directly determines the antenna gain and, for a given operating frequency, the physical size of the antenna aperture (see a further discussion on antennas in Section 2.7). K21636 Book.indb 30 10/22/13 4:09 PM 31 Satellites Satellite h_0 Orbit 2α R_s M θ Coverage area G R_e γ Equator ϕ Figure 2.6 An illustration of coverage area and apex angle. Using the notations in Figure 2.6 as a guide, the coverage area, A_{cov} , from which the satellite is visible with an elevation angle of at least θ can be established as $A_{\text{cov}} = 2\pi R_e^2 (1 - \cos \gamma)$ (2.14) where γ is the central angle. It is a spherical trigonometric relation that relates to the earth and satellite coordinates (to be discussed in detail in Section 2.4). The apex angle required at the satellite to produce a given coverage A_{cov} must satisfy $2\pi R_e^2 (1 - \cos \alpha) = A_{\text{cov}}$ (2.15) However, for small angles, that is, α

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